

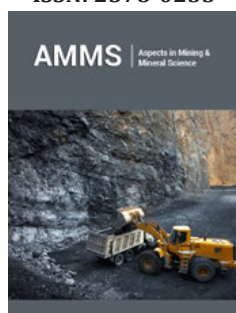
# Thermodynamics of Vacancy Formation in the CoCrFeMnNi High Entropy Alloy from DFT Calculations

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## Abstract

Vacancy formation in High Entropy Alloys (HEA) is an important aspect of their thermodynamic and kinetic theoretical description. Recent studies have shown that this topic can be approached from the side of density functional theory investigations for some concentrated alloys. However, DFT modeling of vacancy formation in HEA remains a challenging and often too complex task due to large sets of atomic configurations near a point defect that need to be accounted for in the calculations. Another challenge is related to an additional degree of complexity related to the paramagnetic state of some HEA such as the classical equimolar fcc CoCrFeMnNi. In this study, we demonstrate a way to reduce the computational complexity of point defect calculations in HEA and investigate the formation of thermal vacancies in the paramagnetic fcc CoCrFeMnNi alloy using DFT-based models and experimental X-ray diffraction analysis of the alloy crystal structure. We evaluate the effect of the paramagnetic state with respect to standard non-magnetic calculations. The calculated results predict a sizable contribution from the magnetic state with yet minor variations of the effective vacancy formation energy in the range from 1.66 at the room temperature up to 1.76 eV at 1373K.

**Keywords:** High entropy alloys; Point defects; Density functional theory; Paramagnetism; Disordered alloys

## Introduction

High Entropy Alloys (HEA) remain in focus of intensive research over the last decade due to their outstanding properties [1]. HEA are often defined as single phase alloys consisting of multiple base elements varying their composition in the range from 5 to 35 at. % [1]. The most known representative of HEA is the equiatomic face centered cubic CoCrFeMnNi - or Cantor - alloy [2]. The name originates from the first idea that the high mixing entropy stabilizes the single-phase solid solution [1,2]. However, nowadays, also multiphase alloys - mostly denoted as Complex Concentrated Alloys (CCA's) - with varying elemental content exhibiting various atomic ordering effects are being considered as HEA as well. As they are envisioned to play an important role in the future of this alloy class [3,4] they are also studied in detail by computational methods [5]. Most of the kinetic processes in HEA are related to vacancy mediated diffusion, which is known to be retarded in some HEA and often referred to as the sluggish diffusion [1]. However, other studies report contradictory results, showing no significant change of the diffusion coefficient compared to the elemental metals [4], calling for detailed investigations of this phenomenon. Vacancy formation and migration processes have also been reported to be an important aspect in development of advanced irradiation resistant alloys for applications in nuclear fusion and -fission reactors [6]. During irradiation, vacancies are formed and actively participate in the main damage mechanisms in these materials [7,8]. Another application area of HEA that draws attention with regard to point defect formation is the production of oxide dispersion strengthened materials for applications at elevated

temperatures such as in energy conversion or propulsion systems [9-11]. During mechanical alloying, the process of incorporating the oxide particles, vacancies seem to foster the dissolution or refinement of oxides leading to enhanced mechanical properties and shorter milling times and therefore a detailed knowledge of vacancy formation processes in these materials is of high interest [10-13].

Investigation of the vacancy formation in compositionally complex HEA using predictive density functional theory methods has drawn much attention in the recent years [14-24]. Compared to point defect formation in one-component materials, the problem of point defect formation in HEA is related to its dependence on the local chemical environment, which makes both experimental and theoretical investigations a challenging task. In the case of theoretical studies, vacancy formation in HEA is strongly connected to sampling of the atomic calculations over a set of crystal structures with various local chemical environments. This task alone represents a challenge for modern predictive density functional theory methods at 0K [25] and the most recent studies suggest employing machine learning approaches to provide a computationally efficient solution to the problem [14,24]. This problem gets even more complicated for magnetic systems, where the chemical disorder is often intertwined with the magnetic disorder that adds an extra degree of complexity to the theoretical description of the material that cannot be routinely modeled by most of the existing DFT codes. Adding the temperature dependence to this task, including both lattice expansion and anharmonic vibrations, that have been shown to be an important aspect of a theoretical evaluation of the vacancy formation, when comparing to experimental data [26-31], seems to be a prohibitively expensive task for most of the HEA systems from the computational point of view and therefore it has not been directly addressed so far to the best of our knowledge. Whereas an efficient general solution to the complete temperature/composition/magnetic state ab initio description of HEA may still require some time to be implemented, a theoretical description of the vacancy formation in paramagnetic HEA can be approached from the side of existing mean-field approaches for the electronic structure and configurational thermodynamics calculations [32]. In Ref. [32], it has been shown that some simple mean-field considerations along with statistical methods can be efficiently used to describe the configurational contribution to the formation energy of vacancies in binary solid solutions at elevated temperatures. A special role in this investigation has been given to the effective vacancy formation energy, which was shown to be significantly lower than the mean value of the local vacancy formation energy calculated in most of DFT investigations and strongly dependent on temperature and configurational effects. In this work, we would like to extend and apply some parts of the aforementioned approach to a multicomponent system and demonstrate an example of the vacancy formation energy calculation in paramagnetic fcc CoCrFeMnNi HEA.

## Methods

### DFT computational details

The nonmagnetic (NM) total energy calculations have been performed by the Projector Augmented Wave (PAW) method [33] implemented in the Vienna Ab Initio Simulation Package (VASP) [34,35]. The Perdew-Burke-Ernzerhof (PBE) form of the Generalized Gradient Approximation (GGA) [36] was used for the exchange-correlation energy. The energy cutoff was 270 eV. The Brillouin zone integration was done using the k-point mesh approximately equivalent to the  $12 \times 12 \times 12$  grid [37] for the 4-atom conventional fcc cubic unit cell. The magnetic contribution has been obtained by the EMT0 (exact muffin-tin orbitals) [38-40]-LSGF [41] (locally self-consistent Green function method) method [42] which allows disordered local moment (DLM) supercell calculations within the Coherent Potential Approximation (CPA) [43,44]. All the self-consistent EMT0-LSGF calculations were performed by using an orbital momentum cutoff of  $l_{\max}=3$  for the partial waves in the spdf basis with the local interaction zone for electronic multiple scattering processes set to three nearest neighbor shells in the supercell calculations. The total energies were obtained using the GGA-PBE functional [36]. The magnetic entropy contribution at elevated temperatures has been included in a model form of longitudinal spin fluctuations (LSF) [45].

### Vacancy formation energy

In contrast to one-element-based systems, vacancy formation in alloys depends on the local atomic environment around a lattice site where it is formed. In a general case, one can speak of the local vacancy formation energy distribution function  $g(E_f)$  that can be defined as

$$\int dE_f g(E_f) = 1 \quad (1)$$

Here, the formation energy of each thermal vacancy,  $E_f$ , is calculated using two types of supercells designed to preserve the chemical composition of the alloy. Here, we have used a  $4 \times 4 \times 4$  supercell made by replication of the 4-atom conventional fcc cell (256 sites) to model a vacancy in the 5-component equimolar CrMnFeCoNi alloy with 20% (51 atom) of each atomic species. Another  $5 \times 5 \times 5$  supercell made by replication of the primitive 1-atom fcc cell (125 sites) has been used to model the bulk reference with ideal lattice (no vacancy). In both supercells, the atoms have been distributed on the lattice to make sure that the pair correlation functions between all atomic species on the first two coordination spheres are as close as possible to the values of a disordered alloy. Due to cell-size limitations, it was impossible to reach identical short range order values for both cell types and the SRO values of the smaller (125 site) cell have been adopted in both calculations. In this case, the vacancy formation energy of an  $i^{\text{th}}$  alloy configuration reads:

$$E_f^i = E_{vac}^i - \frac{a}{b} E^i \quad (2)$$

where  $a$  is the number of atoms in a cell with a vacancy (255) and  $b$  is the number of atoms in a cell without the vacancy (125). The convergence of this configuration with respect to the SRO has been checked by a test calculation of a  $5 \times 5 \times 5$  supercell made by replication of the 4-atom conventional fcc cell (500 sites) with the SRO parameters of the 256-site cell. The vacancy formation energy difference between two supercell setups was less than 0.008 eV. According to Ref. [32], the local vacancy formation energy distribution function  $g(E)$  in an equiatomic binary alloy can be approximated by the discrete binomial distribution. In the case of an equiatomic multicomponent alloy, this distribution has to be replaced by a multinomial distribution function:

$$g(E(m)) = \frac{n!}{m_1! \dots m_k!} c_1^{m_1} \dots c_k^{m_k} \quad (3)$$

where  $k$  is the number of alloy components;  $c_k$  is the atomic fraction of the  $k^{\text{th}}$  component ( $\sum c_k = 1$ );  $n$  is the number of nearest neighbors in the considered coordination shell (12 in the case of the 1<sup>st</sup> nearest neighbor shell in the fcc crystal structure); each  $m = (m_1, \dots, m_k)$  is the number of type  $k$  atoms in the given coordination shell (their sum equals to  $n$ );  $E(m) = (E_1, \dots, E_k)$  are the vacancy formation energies as a function of  $m$  calculated using Eq. 2, which can also be expressed as  $E(m) = E^0 + \sum_k m_k V_k^{-1}$  where  $E^0$  is the lowest local vacancy formation energy corresponding to  $n=0$  and  $V_k^{-1}$  is the interaction energy between the vacancy and  $k^{\text{th}}$  alloy component in the 1<sup>st</sup> coordination shell [32]. The latter can be determined either by the slope of the  $E(m_k)$ , assuming that interatomic interaction in the system is relatively weak, or by the multivariate regression analysis of  $E(m)$ . An equivalent representation of the vacancy formation energy distribution function from (32) is:

$$g(E) = \frac{1}{\sigma\sqrt{2\pi}} \exp \left[ -\frac{(E - \langle E_f \rangle)^2}{2\sigma^2} \right] \quad (4)$$

where  $\sigma$  is its standard deviation,  $E_f$  is the mean value of the vacancy formation energy. This expression can be adopted to obtain the effective local vacancy formation energy as:

$$E_f^{\text{eff}} = \langle E_f \rangle - \frac{\sigma^2}{T} \quad (5)$$

## Experimental details

Investigated powder material was the equiatomic face centered cubic CoCrFeMnNi alloy exhibiting a mean particle size of 400  $\mu\text{m}$ .

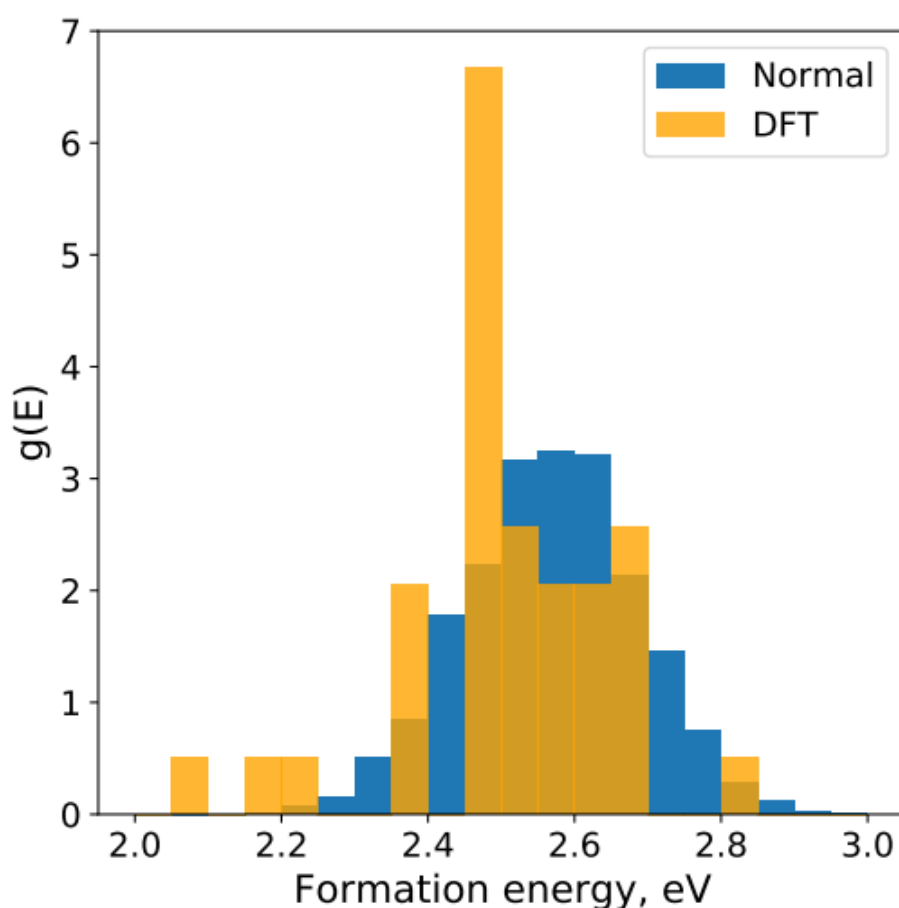
It was a powder milled for 48h under high vacuum in a specifically constructed attritor mill. In situ X-ray diffraction patterns of the equiatomic face centered cubic CoCr-FeMnNi alloy were measured at the HEMS beamline of PETRA III [46] at Deutsches Elektronen-Synchrotron (DESY) in Hamburg using high-energy synchrotron radiation. The powder was put in a Nb-capsule and measured at 25 °C as well as at 1100 °C with a heating rate of 100K/min in-between. For synchrotron measurements a detector distance of 1.4m and a wavelength of 0.14235 Å was used. For evaluation regarding lattice parameter the patterns were fitted using a Pseudo-Voigt model and the peak positions were measured from the maximum of the fit. Due to planar faults present in the investigated material after milling, asymmetric peak shifts occurred and were corrected using the theoretical formula described by Warren [47] and an error minimization approach for all measured peaks. By this, the effect of planar faults on the lattice parameter was corrected.

## Results and Discussion

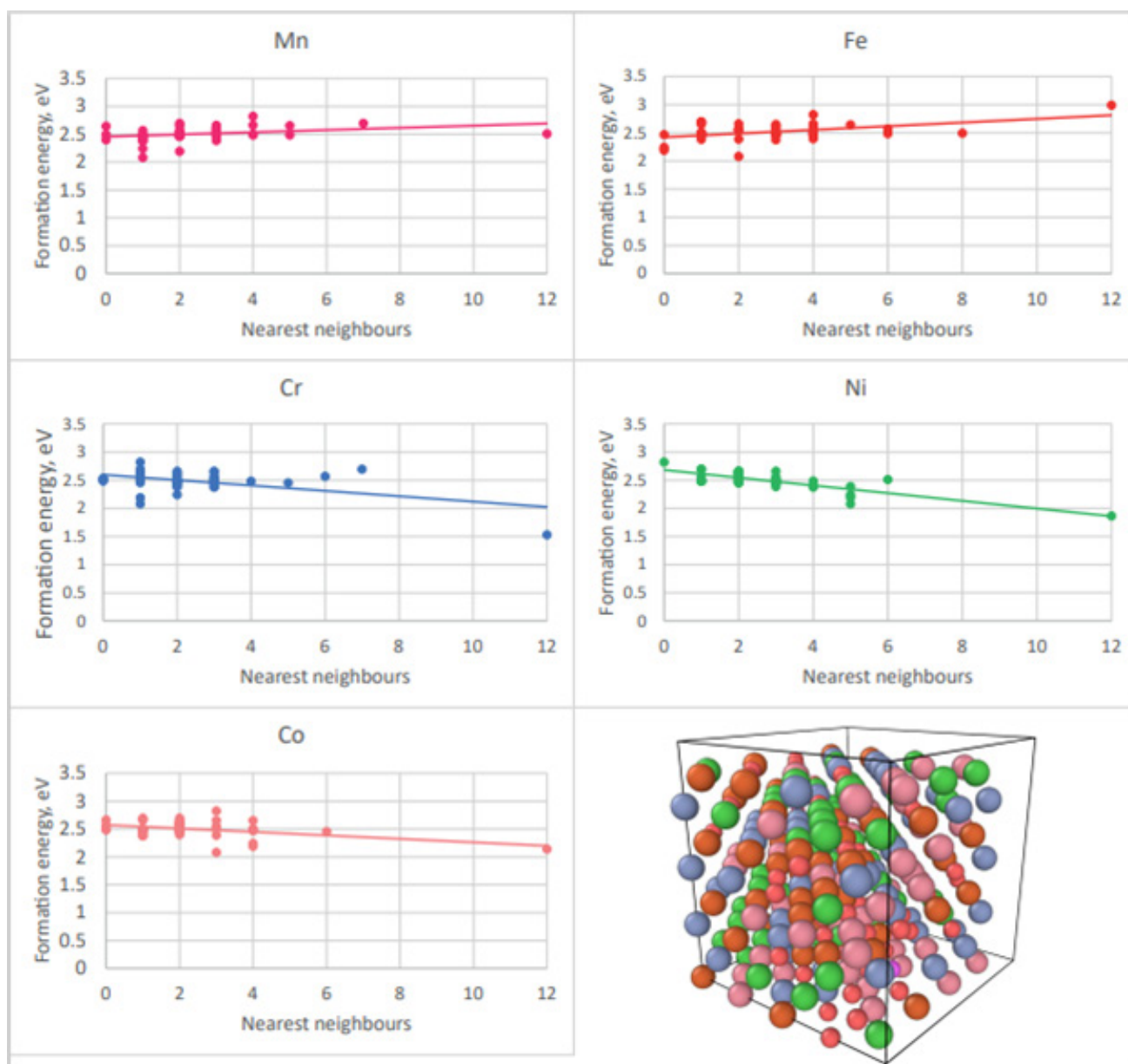
DFT calculations of the vacancy formation in multicomponent alloys is not a trivial task that may become unaffordable expensive as the number of alloy components increases. Currently available technical solutions to this problem usually involve either supercell techniques that rely on averaging over large-enough sets of atomic configurations or on CPA-based methods based on the mean field representation of an alloy. Both methods have their advantages and disadvantages when it comes to the calculation of defects in an alloy. In this paper, we demonstrate a computational approach for DFT calculations of the vacancy formation energy that can combine both methods and benefit from their strong sides. Our computational procedure consists of two main steps: i) calculation of the minimal set of supercells representing the equimolar disordered alloy with a vacancy in the NM state; ii) calculation of the magnetic contribution to the vacancy formation energy using DLM-CPA. Here, we focus on formation of thermal vacancies which are formed without changing the chemical composition of the alloy. Therefore, we have used two different supercells to represent the system with and without a vacancy chosen in a way that they both have the same equimolar chemical composition (of 20 at. % for each atomic species) as described in Section 2.2. In (Figures 1 & 2), we show the local vacancy formation energy distribution function and the formation energy dependence on the number of nearest neighbor atoms of each individual alloy component, respectively. These results have been obtained in the NM calculations using the Room Temperature (RT) lattice parameters obtained from the XRD measurements (Table 1). Figure 1 also shows the scatter of the vacancy formation energies that covers the range between 2 and 3 eV.

**Table 1:** Alloy lattice parameters ( $a$  in Å), vacancy formation energies (in eV) with corresponding temperatures (in K) and the standard deviations of the local vacancy formation energies using Eq. 3 fit to DFT data. The correction from paramagnetic state from LSF and DLM is denoted by  $\Delta(E_f)$ .

| Reference   | $a$    | $\langle E_f \rangle$ | $\Delta(E_f)$ | $E_f^{\text{eff}}$ | T         | $\sigma$ |
|-------------|--------|-----------------------|---------------|--------------------|-----------|----------|
| Exp.        | 3.5928 | -                     | -             | -                  | 298       | -        |
| DFT+RTL+LSF | 3.5928 | 2.13                  | 0.32          | 1.66               | 298       | 0.118    |
| Exp.        | 3.6740 | -                     | -             | -                  | 1373      | -        |
| Exp. (22)   | -      | -                     | -             | 1.69               | 1181-1423 | -        |
| DFT+RTL+LSF | 3.6740 | 1.85                  | 0.53          | 1.76               | 1373      | 0.120    |



**Figure 1:** Local vacancy formation energy distribution function  $g(E)$  obtained from DFT calculations (orange) and approximated by Eq. 4 (blue).



**Figure 2:** Local vacancy formation energy dependence on a number of nearest neighbor atoms of alloy components in the supercell. Linearly extrapolated data are shown with solid lines. Symbol and line coloring in the graphs corresponds to the atom coloring in a schematic representation of the supercell used in the calculations.

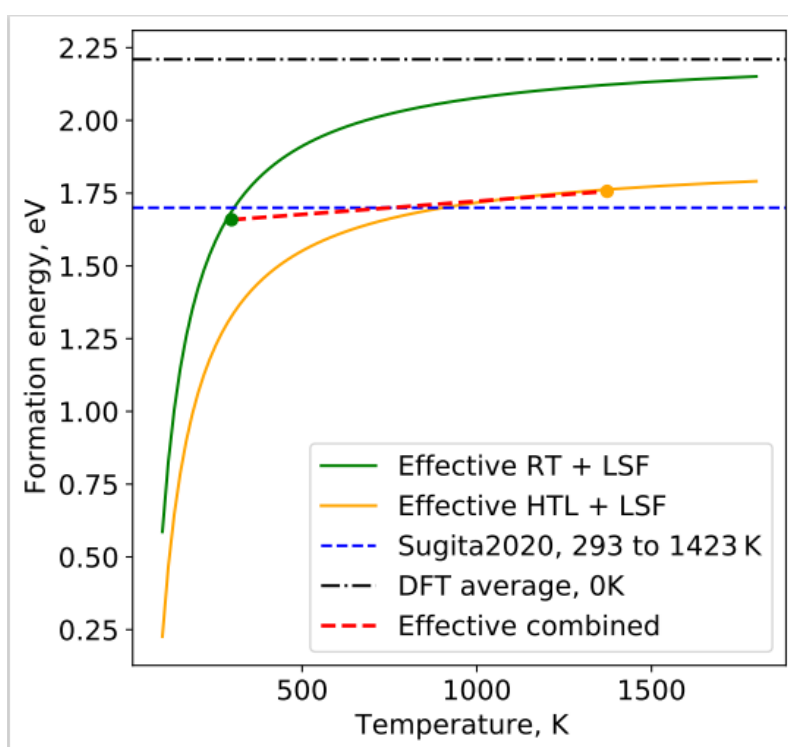
A more detailed analysis of the distribution of the local vacancy formation energies with respect to the number of nearest neighbor atoms of different sorts can be seen in Figure 2. Here, by analogy with binary alloys [32], we analyze individual contributions of each alloy component to the vacancy formation energy in the alloy. The slopes of the linear fits to these data correspond to the interaction energy of a vacancy with the atoms of each selected alloying elements, assuming that interactions between the atoms themselves are relatively weak. The number on the x-axis in Figure 2 corresponds to the number of atoms in the first coordination shell varying from 1 to 12. There has been no site surrounded by 12 nearest neighbor

atoms of the same species in the analyzed supercells and we have used a set of additional calculations where the cell volume and fcc structure were kept fixed, but all atoms were replaced solely by one type of atomic species to mimic the limiting case of 12 nearest neighbor atoms of the same sort. The results of this investigation suggest that an increase of the number of Cr, Ni and Co atoms at the first coordination shell around the vacancy reduces the  $E_f$  on average, whereas Mn and Fe slightly increase it. This result can also be interpreted in a way that the vacancy is likely to be surrounded by the atoms of these sorts in the alloy. Mn has virtually no effect on the  $E_f$  and the strongest effect belongs to Ni that can lower the



$E_i$  by up to 20% in our estimate of the limiting case of its 100% occupation of the 1<sup>st</sup> coordination shell and only by around 10% at the 20% concentration corresponding to its content in the Cantor alloy. These results also indicate that the interatomic interactions in the considered alloy are relatively small compared to the formation energy of a vacancy itself. The results on the vacancy interaction energies with alloying elements allow us to reconstruct the normal multinomial distribution function using Eq. 3. Assuming that shape of this distribution function is generic for concentrated equimolar alloys [32], one can use it to obtain an accurate (physical) fit of the calculated in DFT data and extract necessary for calculating  $E_i^{\text{eff}}$  standard deviation and mean formation energy values (Table 1). The result of the fit is shown in (Figure 1) in blue and indicates

some differences between the normal and DFT-based distributions, which are most likely related to the limited set of selected atomic configurations for DFT calculations. However, we observe a rather fast convergence of the results of the fit even if one uses a very limited set of DFT data: Already a set of 11 atomic configurations selected according to Sec. 2.2 is sufficient to get the mean formation energy value within 0.012 eV and the  $\sigma$  within 0.019 error and the results for >40 alloy configurations calculated in DFT lead to practically no change in the fit parameters. This way, we see a great advantage in the use of Eq. 3 for analysis of the vacancy formation in equimolar alloys as it can substantially reduce the computational cost related to the DFT calculations of various defect configurations.



**Figure 3:** Effective vacancy formation energy as a function of temperature, calculated at the RT and HTL DFT input data (solid lines). Vertical dashed lines represent the experimental from Ref. (22) (blue) and the mean 0K DFT values (black). Two circles connected with a red dashed line represent the results of the effective vacancy formation energy predictions based on both DFT and model calculations done at the corresponding temperatures (298 and 1373K) suitable for comparison to the experimental data.

The calculated values so far have been obtained in the nonmagnetic state approximation. However, the alloy under consideration is paramagnetic and therefore we have applied the DLM and LSF models to calculate the contribution of paramagnetism to the calculated vacancy formation energies in the NM state. The calculations have been done at two representative temperatures denoted as RT (room temperature, 298K) and HT (high temperature, 1373K) at the corresponding lattice parameters (RTL and HTL respectively) measured in XRD (Table 1). The results show that the NM calculations overestimate the  $E_i$  by 0.32 and 0.53 eV (Table 1) compared to the DLM RTL+LSF and HTL+LSF calculations. We

have applied this correction to our NM calculations to calculate the effective local vacancy formation energy  $E_i^{\text{eff}}$  using Equation 5 and to compare it to the experiment. An important point in this analysis is that the effective local vacancy formation energy is temperature dependent and that it is always (at  $T < \infty$ ) smaller than the mean vacancy formation energy value,  $\langle E_i \rangle$  (see Equation 5). These results can be clearly seen in (Figure 3), where  $E_i^{\text{eff}}$  is plotted along with the mean  $\langle E_i \rangle$  and experimental vacancy formation energy values, as well as in Table 1. The  $E_i^{\text{eff}}$  values presented in the table for RT and HTL calculations yield 1.66 and 1.76 eV respectively. These results are very close to the experimental value of 1.69 eV

[22], which refers to the positron lifetime measurements. In Ref. [22], the authors were able to single out the formation of thermal vacancies that were found to be dominant in the CoCrFeMnNi HEA in the temperature interval from 1181 to 1423K and declared that the effective formation energy of them remain virtually a constant within this interval. If we connect two calculated points referring to RT and HTL in (Figure 3), we see that this conclusion is consistent with the theoretical prediction of this work, where we can see only a minor change in  $E_f^{\text{eff}}$  as a function of temperature. Here, we must note that Equation 5 is strictly speaking valid in the high temperature regime, where the  $E_f^{\text{eff}}$  changes are indeed linear. In the limit of  $T \rightarrow 0$ , the strong deviations from linearity are related to its analytical form, which leads to most likely artificial reduction of  $E_f^{\text{eff}}$  and therefore underestimation of the  $E_f^{\text{eff}}$  at RT in our calculations. Another important aspect is the partially neglected effect of the atomic vibrations that may have an impact on the resulting  $E_f^{\text{eff}}$  values at high temperatures [28-31]. These effects in the selected alloy need to be investigated further in more detail, though they may have their main contribution to the vacancy formation in the considered system through the lattice expansion that is included in the calculation via adaptation of the experimental lattice constants. The remaining (not included in this work) part of it may have a minor impact on the thermal vacancy formation in the considered alloy, as we can see from the comparison of our results to the experimental data.

## Conclusion

Vacancy formation in multicomponent paramagnetic alloys is a complex topic for density functional theory investigations. However, the complexity of the problem can be reduced by the use of a number of model steps that can make such calculations a feasible task. In this paper, we demonstrate the applicability of existing theoretical approaches for DFT-based calculations of the effective vacancy formation energy calculations and extend them to the paramagnetic fcc CoCrFeMnNi HEA. We perform two sets of DFT-based calculations corresponding to the room and high temperature material states and analyze the obtained vacancy formation energies as a function of thermal lattice expansion and the magnetic state. We show that the account of the magnetic degrees of freedom in the calculations leads to a sizable contribution to the formation energy of the defects. Our results show that the formation energy of thermal vacancies in the paramagnetic fcc CoCrFeMnNi HEA varies little as a function of temperature taking values from 1.66 to 1.76 eV between 298 and 1373K, which agrees well with available experimental data.

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